

1. A satellite in a circular orbit

A satellite, mass m_1 , orbits a planet of mass m_2 at distance r from the planet's center. $m_1 \ll m_2$

A. The info given suffices to give unique values for V , U_g , L of satellite

TRUE $U_g = -\frac{Gm_1m_2}{r}$, $L = m_1vR$, $v = \sqrt{\frac{GM_2}{r}}$

B. The total mechanical energy is conserved

TRUE No dissipative forces act on the system

C. The linear momentum vector of the satellite is conserved.

FALSE Since \vec{v} changes direction, $\vec{p} = m\vec{v}$ is always changing as well.

This is because there is a nonzero net force on the satellite, & N2L says $\vec{F}_{net} = \frac{d\vec{p}}{dt} \neq 0$

D. \vec{L} of satellite about planet's center is conserved

TRUE Because \vec{F}_g is directed toward the center, its torque is zero ($\vec{\tau} = \vec{r} \times \vec{F}$, but \vec{r} & \vec{F} are 180° apart, so $|\vec{r} \times \vec{F}| = rF \sin \theta = rF \sin 180^\circ = 0$)

Since $\vec{\tau}_{net} = \frac{d\vec{L}}{dt}$, \vec{L} is constant $\leftrightarrow \vec{L}$ is conserved

E. One can calculate orbital speed from conservation of E & \vec{p}

FALSE For one thing, \vec{p} is not conserved (see C.)

2. Ex 8.30 $v_{esc} = 7.5 \frac{\text{km}}{\text{s}}$ for a planet whose mass is $3.3 \times 10^{24} \text{ kg}$. What is the planet's radius?

Since $v_{esc} = \sqrt{\frac{2GM}{r}}$, $v_{esc}^2 = \frac{2GM}{r} \rightarrow r = \frac{2GM}{v_{esc}^2} = \frac{2(6.67 \times 10^{-11} \text{ N} \frac{\text{m}^2}{\text{kg}^2})(3.3 \times 10^{24} \text{ kg})}{(7.5 \times 10^3 \frac{\text{m}}{\text{s}})^2}$

$r = 7.83 \times 10^6 \text{ m} \rightarrow \boxed{7.8 \times 10^6 \text{ m}}$

3. Two identical spheres, $m = 7.75 \text{ kg}$, experience $F_g = 0.285 \mu\text{N}$. Alt Ex 8.77
 What is the separation between their centers?

$$F_g = \frac{G m_1 m_2}{r^2} \rightarrow r^2 = \frac{G m_1 m_2}{F_g} \rightarrow r = \sqrt{\frac{G m_1 m_2}{F_g}} = \sqrt{\frac{6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2} (7.75 \text{ kg})^2}{0.285 \times 10^{-6} \text{ N}}}$$

$$r = 0.119 \text{ m} \quad \text{or } 11.9 \text{ cm}$$

4. CQ 8.01 A baseball is at Earth's surface. True statements include:
 The grav. force of the ball on Earth is exactly the same as the force by Earth on ball (in magnitude). True by N3L

5. CQ 8.04

Suppose Earth (& everything on it) suddenly doubled in density (mass doubles, size does not change). How would g change?

Since $g = \frac{GM}{r^2}$, doubling M (only) doubles g

6. Orbital speed of a satellite

A. Satellite 1: speed = v , orbit radius r

satellite 2: ~~v~~ speed $< v$ How does its orbital radius compare to r ?

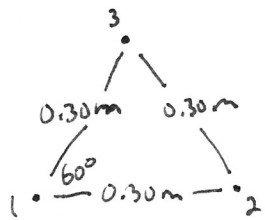
Since $v = \sqrt{\frac{GM}{r}}$, smaller v implies larger $r \rightarrow$ Distance $> r$

B. Satellite 1 orbits Earth at r, v_E } compare T_E to T_M
 " " Moon at r, v_M

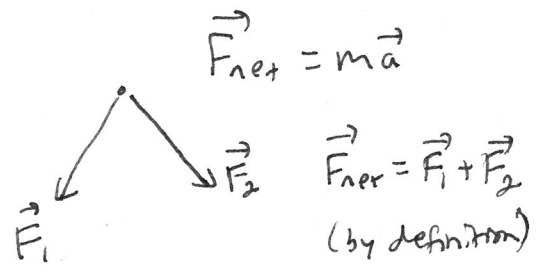
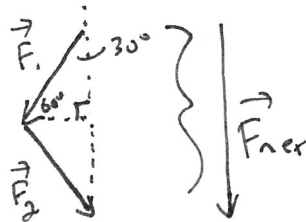
Again, since $v = \sqrt{\frac{GM}{r}}$ and $M_E > M_M$, the satellite in Earth orbit has greater speed & thus shorter period. So $T_E < T_M \rightarrow$ $T_M > T_E$

7. Prob. 8.03

3 identical 50 kg masses are at vertices of equilateral triangle. one is released from rest. What is its initial acceleration?



Let's release #3. FBD:
Apply N2L



The x components of \vec{F}_1 & \vec{F}_2 are going to cancel by symmetry. Meanwhile

the magnitude of $F_{net} = |F_{1y}| + |F_{2y}|$. Also note that $F_1 = F_2 = F_g$ - both individual vectors will have the same magnitude, since all masses and distances are equal. Let's get the y component of one of them

$$F_{1y} = F_1 \cos 30^\circ = \frac{G m_1 m_3}{r^2} \cos 30^\circ = \frac{G m^2}{r^2} \cos 30^\circ \quad (\text{since } m_1 = m_3 = m = m_2)$$

$$\vec{F}_{net} = (F_{1y}) + (F_{2y}) = 2|F_{1y}| = 2 \frac{G m^2}{r^2} \cos 30^\circ$$

$$\text{Since } F_{net} = ma, \quad a = \frac{F_{net}}{m} = \frac{1}{m} 2 \frac{G m^2}{r^2} \cos 30^\circ = \frac{2GM}{r^2} \cos 30^\circ$$

$$a = \frac{2 (6.67 \times 10^{-11} \text{ N } \frac{\text{m}^2}{\text{kg}^2}) (50 \text{ kg})}{(0.3 \text{ m})^2} \frac{\sqrt{3}}{2} = \boxed{6.4 \times 10^{-8} \frac{\text{m}}{\text{s}^2}}$$

8. Problem 8.07 By how many Newtons does the weight of a $m=100\text{ kg}$ person change going from sea level to 5.0 km ?

$$F_G = \frac{GM_E m}{r^2} \quad \Delta F_G = \frac{GMm}{\left(\frac{1}{r_{E+5\text{ km}}}\right)^2} - \frac{1}{r_E^2}$$

$$\Delta F_G = (6.67 \times 10^{-11} \text{ N} \frac{\text{m}^2}{\text{kg}^2}) (5.97 \times 10^{24} \text{ kg}) (100 \text{ kg}) \left(\frac{1}{(6.37 \times 10^6 + 5 \times 10^3 \text{ m})^2} - \frac{1}{(6.37 \times 10^6 \text{ m})^2} \right)$$

$$\Delta F_g = -1.5 \text{ N}$$

Change should be negative, and a small fraction of weight (since 5 km is a small fraction of 6400 km)

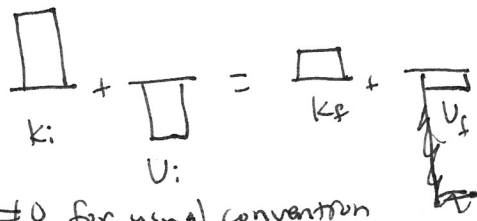
9. Problem 8.15

Planet w/ $M = 2.92 \times 10^{23} \text{ kg}$, $R = 5.00 \times 10^6 \text{ m}$. Projectile fired straight up at 2000 m/s .
What is speed when $h = 1000 \text{ km}$?

Apply conservation of mechanical energy

$$K = \frac{1}{2}mv^2, \quad U_g = -\frac{GMm}{r}$$

Note that $U_i \neq 0$ for usual convention



$$\frac{1}{2} m v_i^2 - \frac{GMm}{R} = \frac{1}{2} m v_f^2 - \frac{GMm}{(R+1000\text{ km})}$$

\uparrow \uparrow \uparrow \uparrow
 K_i U_i K_f U_f

Multiply through by $\frac{2}{m} \dots$

~~$$v_f^2 = v_i^2 + 2GM \left(\frac{1}{R+h} - \frac{1}{R} \right)$$~~

$$v_f = \sqrt{v_i^2 + 2GM \left(\frac{1}{R+h} - \frac{1}{R} \right)} = \sqrt{(2000 \frac{\text{m}}{\text{s}})^2 + 2(6.67 \times 10^{-11} \text{ N} \frac{\text{m}^2}{\text{kg}^2})(2.92 \times 10^{23} \text{ kg}) \left[\frac{1}{6 \times 10^6 \text{ m}} - \frac{1}{5 \times 10^6 \text{ m}} \right]}$$

$$= 1.64 \times 10^3 \frac{\text{m}}{\text{s}}$$

$$v_f = 1640 \frac{\text{m}}{\text{s}}$$

10. Problem 8.18

A 910 kg object is released from $h = 1200$ km. What is speed when it strikes Earth (ignoring air resistance)?

Just like # 9... $k_i = 0$

$$\overline{k_i} + \overline{k_{U_i}} = \overline{k_f} + \overline{k_{U_f}}$$

$$0 - \frac{GMm}{(R_E + h)} = \frac{1}{2}mV^2 - \frac{GMm}{R_E}$$

$$V^2 = 2GM \left(\frac{1}{R_E} - \frac{1}{R_E + h} \right)$$

↓

$$V = \sqrt{2GM \left(\frac{1}{R_E} - \frac{1}{R_E + h} \right)}$$

$$= \sqrt{2 (6.67 \times 10^{-11} \text{ N m}^2 / \text{kg}^2) (5.97 \times 10^{24} \text{ kg}) \left[\frac{1}{6.37 \times 10^6 \text{ m}} - \frac{1}{(6.37 \times 10^6 \text{ m} + 1.2 \times 10^6 \text{ m})} \right]}$$

$$= \sqrt{.98 \times 10^7 \frac{\text{m}^2}{\text{s}^2}} = 4.45 \times 10^3 \frac{\text{m}}{\text{s}} \rightarrow \boxed{4.5 \frac{\text{km}}{\text{s}}}$$

11. Problem 8.34

A high jumper can jump ~~2.0~~ 2.45 m on Earth. How high would he jump on

A. Mars ($g_M = 3.74 \frac{m}{s^2}$)

We'll assume he does the same work in jumping, which will convert to gravitational potential energy change. Then

$$m g_E h_E = m g_{Mars} h_{Mars}$$

$$h_{Mars} = h_E \frac{g_E}{g_{Mars}} = 2.45 m \left(\frac{9.81 m/s^2}{3.74 m/s^2} \right) = \boxed{6.43 m}$$

B. Moon ($g_M = 1.62 m/s^2$)

As above, only ~~using~~ using g_{Moon}

$$h_{Moon} = 2.45 m \left(\frac{9.81 m/s^2}{1.62 m/s^2} \right) = 1.48 \times 10^1 m = \boxed{14.8 m}$$